

# Ion Waves → Ion Sound

Plasma → no neutrals and few collisions

In the absence of collisions → ordinary sound waves would not occur.

Ions can still transmit vibrations to each other because of their charge, acoustic waves can occur through the intermediary of an electric field

Massive ions involved → low-frequency oscillations.

We assume  $n_i = n_e = n$  and do not use Poisson's eq<sup>n</sup>. The ion fluid eq<sup>n</sup> in the absence of a magnetic field is

$$Mn \left[ \frac{\partial v_i}{\partial t} + (v_i \cdot \nabla) v_i \right] = enE - \nabla p$$

$$= -en \nabla \phi - \nabla_i kT_i \nabla n$$

— (6)

assumed  $E = -\nabla \phi$  and used eq<sup>n</sup> of state.

Linearizing and assuming plane waves, we have

$$-i\omega Mn_0 v_{i1} = -en_0 ik\phi_1 - \nabla_i kT_i ikn_1$$

— (7)

The balance of forces on electrons, requires

$$n_e = n = n_0 \exp\left(\frac{e\phi_1}{kT_e}\right) = n_0 \left(1 + \frac{e\phi_1}{kT_e} + \dots\right)$$

The perturbation in density of electrons, and, therefore, of ions, is then

$$n_1 = n_0 \frac{e\phi_1}{kT_e} \quad \text{— (8)}$$

Incorporated in eq<sup>n</sup> of continuity

$$i\omega n_i = n_{i0} ik v_{i1} \quad \text{--- (9)}$$

In eq<sup>n</sup> (7), we may substitute  $\phi_1$  and  $n_i$  in terms of

$v_{i1}$  from eq<sup>n</sup> (6) & (9)

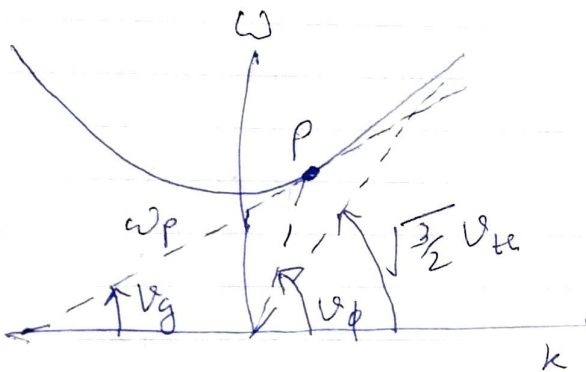
$$i\omega M n_0 v_{i1} = \left( e n_0 ik \frac{k T_e}{e n_0} + \sum_i \gamma_i k T_i ik \right) \frac{n_0 (ik v_{i1})}{i\omega}$$

$$\omega^2 = k^2 \left( \frac{k T_e}{m} + \sum_i \frac{\gamma_i k T_i}{m} \right)$$

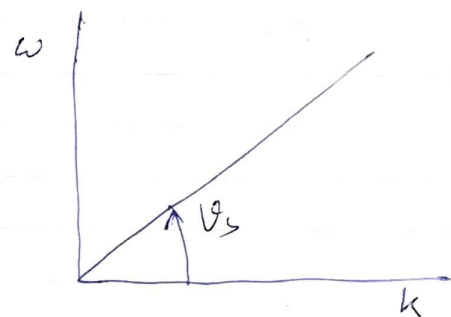
$$\frac{\omega}{k} = \left( \frac{k T_e + \sum_i \gamma_i k T_i}{m} \right)^{1/2} \equiv v_s \quad \text{--- (10)}$$



Sound speed in Plasma



Dispersion relation for  
electron plasma waves  
(Bohm-Gross waves)



Dispersion relation for  
in acoustic waves in  
the limit of small  
Debye length.

Plasma oscillations → basically constant-frequency waves

with a correction due to thermal motions

Ion waves → constant-~~velocity~~ velocity waves and exist  
only when there are thermal motions

for ion waves, group velocity is equal to  
the phase velocity.

# Applications of Plasma Physics

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Plasmas → characterized by the two parameters  $n$  and  $kT_e$ , applications cover extremely wide range of  $n$  and  $kT_e$ .

$n$  varies → 28 orders of magnitude from  $10^6$  to  $10^{34} \text{ m}^{-3}$ .

$kT$  → can vary over seven orders from 0.1 to  $10^6$  eV.

## Gas Discharges (Gaseous Electronics)

Earliest work 1920's → Langmuir, Tonks and collaborators.

Vacuum tubes that could carry large currents and therefore had to be filled with ionized gases

Weakly ionized glow discharges and positive columns,  $kT_e \approx 2 \text{ eV}$  and  $10^{14} < n < 10^{18} \text{ m}^{-3}$ .

Spark gaps, welding arcs, neon and fluorescent lights, lighting discharges.

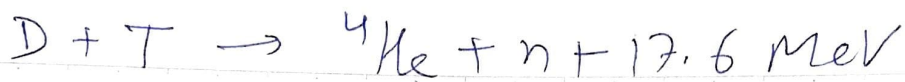
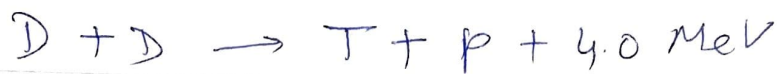
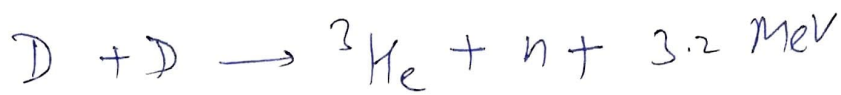
## Controlled Thermonuclear Plasma

1952 → hydrogen bomb fusion reaction

be controlled to make a reactor

The principal reactions involve Deuterium (D) and Tritium (T) atoms





Cross sections for these fusion reactions are appreciable only for incident energies above 5 keV.

Accelerated beams of deuterons bombarding  $\rightarrow$  will not work  $\rightarrow$  deuterons lose energy by scattering  $\rightarrow$  No fusion reaction X

Need of Plasma  $\rightarrow$  Thermal energies in the range of 10-keV.

$\rightarrow$  Heating & containing such a plasma has rapid growth.  $\downarrow$  Confinement

Space Physics  $\rightarrow$  Study of the earth's environment in space.

Solar Wind: Continuous stream of charged particles  $\rightarrow$  impinges on the earth's magnetosphere

$\rightarrow$   $n = 5 \times 10^6 \text{ m}^{-3}$ ,  $kT_i = 10 \text{ eV}$ ,  $kT_e = 50 \text{ eV}$ ,  $B = 5 \times 10^{-9} \text{ T}$   
and drift velocity  $200 \text{ km/sec}$ .

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Ionosphere → from an altitude of 50 km to 10  
earth radii → populated by weakly ionized plasma  
density varying up to  $n = 10^{12} \text{ m}^{-3}$ .  $T$  is only  
 $10^{-1} \text{ eV}$ .

Modern Astrophysics :- Stellar interiors and  
atmospheres are hot enough to be in the  
plasma state.

Core of the sun → 2 keV (Temp)  
Thermonuclear reactions → responsible for  
the sun's radiation.

Solar Corona → Plasma with temp. up to 1000 eV

Plasma as a conducting fluid

- Magnetohydrodynamics (MHD)

Macroscopic Approach:- Plasma is considered  
to be divided into small volumes

Each of which is large compared to the  
average spacing of the individual particles it

contains

Yet small compared to any distance over  
which the macroscopic properties vary  
appreciably.

Associated with these volume

average velocity, magnetic field, density, temperature, pressure, conductivity etc.

→ Behaviour of plasma deduced in terms of the action of these volumes.

Conducting fluid → Plasma → subject simultaneously to the laws of electromagnetism and hydrodynamics

→ Magnetohydrodynamics (MHD) or hydromagnetics.

Concept of magnetic pressure

$$p^m = \frac{B^2}{2\mu_0}$$

introduced by Faraday. He thought of tubes of forces as elastic filaments under tension in the direction of the field and compressed in the transverse direction

Maxwell → stress tensor (mathematical form)

→ Basis for magnetohydrodynamics  
Behaviour of conducting fluid in an electromagnetic field